

Study of Magnetic Island Using a 3D MHD Equilibrium Calculation Code^{*)}

Yasuhiro SUZUKI^{1,2)}, Satoru SAKAKIBARA^{1,2)}, Kiyomasa WATANABE^{1,2)},
Yoshiro NARUSHIMA^{1,2)}, Satoshi OHDAKI^{1,2)}, Satoshi YAMAMOTO³⁾,
Hiroyuki OKADA³⁾ and LHD experiment group

¹⁾National Institute for Fusion Science, Toki 509-5292, Japan

²⁾Graduate University for Advanced Studies (SOKENDAI), Toki 509-5292, Japan

³⁾Institute of Advanced Energy, Kyoto University, Uji 611-0011, Japan

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Coupling the magnetic diagnostics and a 3D MHD equilibrium calculation code, the magnetic island is studied in the Large Helical Device (LHD) experiment. In an experiment, the collapse in the plasma core was observed in a configuration, which has large magnetic island produced by external perturbation coils. At the collapse, the temperature profile was flattened. This suggests the magnetic island evolved. The magnetic island was observed by the magnetic diagnostics. The magnetic diagnostics also suggests evolving the magnetic island. A 3D MHD equilibrium is calculated by the 3D MHD equilibrium code then signals of the magnetic diagnostics are simulated. Since the comparison of observed and calculated signals is comparable, the magnetic island in calculated equilibrium is similar to one of the experiment.

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1. Introduction

The study of the magnetic island is an important issue because the generation of the magnetic island leads the degradation of the confinement due to flattening of the temperature. In tokamaks, the magnetic island is generated by the resistive MHD instability like the tearing mode [1] at many cases. Especially, the study of the neoclassical tearing mode (NTM) [2] is an important and critical issue. On the other hand, in stellarator/heliotron plasmas, the generation of the magnetic island is observed due to the MHD instability [3,4]. In addition, the evolution and suppression of the magnetic island are also observed without the MHD instability [5]. In such cases, the magnetic island is driven by the equilibrium response [6–8]. Since the magnetic island driven by the equilibrium response is not rotating and appeared in the quasi steady-state, it is good target to measure the magnetic island.

To observe the magnetic island, the profile measurement is widely used. Flattening of the electron temperature and density indicate the existence of the magnetic island. However, if the O-point of magnetic islands does not locate on the line of sight, the profile measurement cannot identify the magnetic island. On the other hand, the magnetic diagnostics directly observe the plasma response and it does not depend on the location of the O-point. The magnetic diagnostics observes total plasma response in the out-

side of the plasma. Thus, to identify modes of the plasma response, the mode analysis is done. In the LHD experiments, the mode analysis assuming the current filament is used [9] but this analysis can not model the magnetic field structure. To model the magnetic field more physically, coupling with the numerical simulation is necessary.

An advantage of the LHD device to study the magnetic island is superposing of the resonant perturbation field (RMP) by external coils. These coils are called to the LID coils, which were prepared for the operation of the Local Island Divertor (LID) [10]. Since these coils can generate the low- n magnetic island for the vacuum, we can study only the effect of the plasma response on the magnetic island. Many experiments were done to study the plasma response [5, 11–13]. In those studies, the spontaneous evolution and suppression of the island were observed without the MHD instability. This suggests a possibility the magnetic island changes spontaneously due to the equilibrium response. This also suggests 3D MHD equilibrium analysis can be used to study the magnetic island.

In this study, we study the magnetic island in a low magnetic shear configuration in the LHD. We propose studies of the magnetic island by coupling the magnetic diagnostics and a 3D MHD equilibrium calculation code without assumption of nested flux surfaces. In the next section, we show an experimental result, which is an observation of the perturbation driven by the plasma response. The perturbed field is observed by the magnetic diagnos-

author's e-mail: suzuki.yasuhiro@LHD.nifs.ac.jp

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tics. In Sec. 3, we show a demonstration of our method by coupling the magnetic diagnostics and the 3D MHD equilibrium code. Finally, we discuss and conclude this study.

2. Experimental Results

The LHD is an $L/M = 2/10$ heliotron. Here L and M are the pole number of helical coil winding and toroidal field period. In this study, we did an experiment with $n/m = 1/1$ magnetic island, where n and m are toroidal and poloidal mode numbers, respectively. As mentioned in the introduction, the LHD device can produce low- n magnetic islands by the LID coils. The LID coils produce the dipole field to produce $n/m = 1/1$ field. The plasma response on low- n magnetic islands is observed by the magnetic diagnostics, which are two poloidal arrays of flux loops, “Loop1” and “Loop2”. Figure 1 shows a schematic view of poloidal arrays of flux loops. One array consists of twelve flux loops. These arrays measure the perturbed field produced by the parallel current j_{\parallel} . In the LHD experiments, j_{\parallel} is the Pfirsch-Schlüter (P-S) current. If there is no low- n perturbed field smaller than the toroidal field period, two arrays observe same signals. However, appearing low- n perturbation ($n < 10$), the difference between two arrays appears. From the difference, we can decide the perturbed field B_r , mode numbers n/m and phase of the magnetic island ϕ_{island} .

Figure 2 shows a discharge in a large plasma aspect ratio configuration ($A_p = 8.3$). In this experiments, the

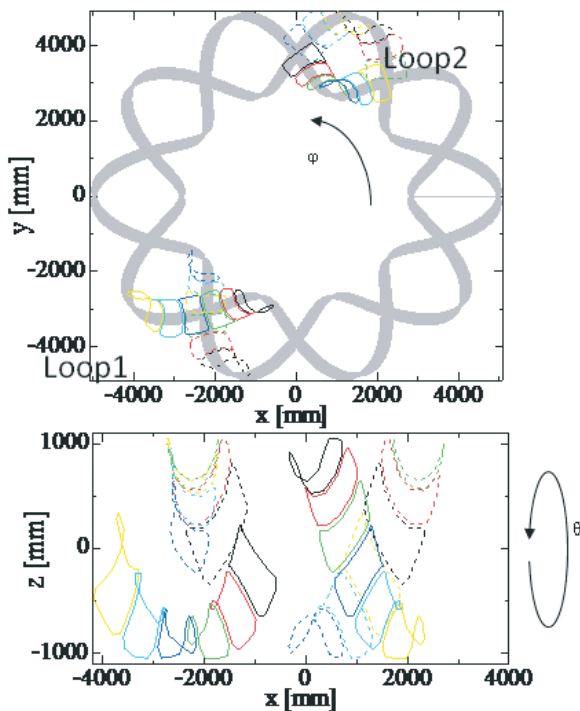


Fig. 1 The schematic view of poloidal arrays of flux loops. Two arrays are installed at $\phi = 57$ and 237 [deg]. One array consists of twelve loops.

dipole field by LID coils was superposed to a configuration, which is $R_{\text{ax}} = 3.6$ m and $\kappa \sim 1$. The R_{ax} is the vacuum axis position and κ is the averaged elongation of the plasma. The dipole field resonates on $\iota = 1$ surface, where ι is the rotational transform. The order of the perturbed field B_{11} is about $O(10^{-4})$. The width of the island is about $\Delta\rho \sim 0.4$, where ρ is the normalized minor radius. For $A_p = 8.3$ configuration, the rotational transform on the axis ι_0 is larger than $n/m = 2/3$ and the rotational transform at the plasma edge on ι_a is smaller than $n/m = 3/2$. This means only $n/m = 1/1$ island appears in this configuration and it can be considered. The magnetic shear on $\iota = 1$ surface is weaker than other configurations with small A_p . Thus, we can expect the equilibrium response will appear strongly. In the figure, the volume averaged beta $\langle\beta\rangle_{\text{dia}}$, the plasma current I_p/B_t , signals Φ_r at a poloidal angle and the difference $\Delta\Phi_r$ between two arrays at the poloidal angle are plotted. A red line indicates the time appearing the minor collapse. Increasing $\Delta\Phi_r$, $\langle\beta\rangle_{\text{dia}}$ slightly decreases. In Fig. 3, profiles of the electron temperature T_e between the

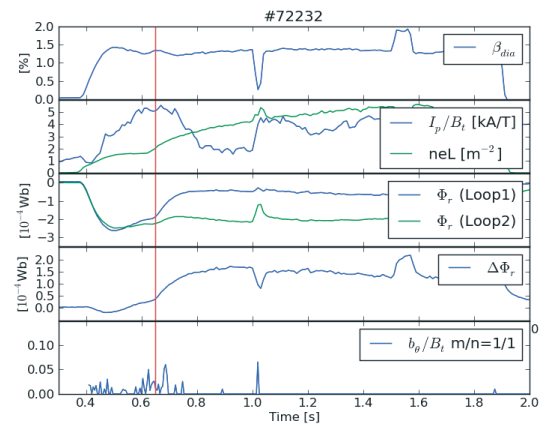


Fig. 2 The volume averaged beta $\langle\beta\rangle_{\text{dia}}$, the plasma current I_p/B_t , signals Φ_r at a poloidal angle and the difference $\Delta\Phi_r$ between two arrays at the poloidal angle are plotted. The minor collapse appears at the red line in the figure.

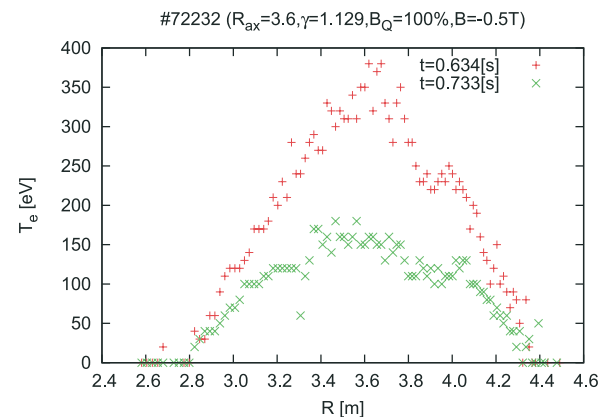


Fig. 3 Profiles of the electron temperature T_e are shown between the minor collapse.

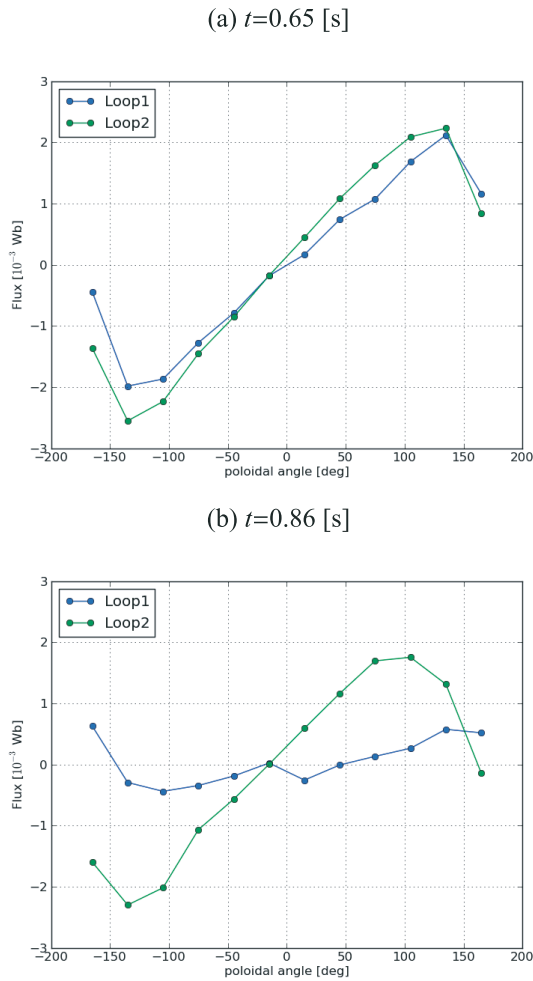


Fig. 4 Measured flux from two poloidal arrays are shown between the collapse $t = 0.65$ and 0.86 [s]. Profiles are plotted along the poloidal angle θ . Differences between Loop 1 and 2 are increased after the collapse.

collapse are plotted. After the collapse, T_e decreases then flattening T_e increases. This suggests the magnetic island with $n/m = 1/1$ evolves.

In Fig. 4, profiles of measured flux on loop arrays are plotted along the poloidal angle θ . Profiles are shown at $t = 0.7$ (before the collapse) and 0.86 (after the collapse). After the collapse, the difference increases and signals of the Loop1 is sensitive. Although the signal of the Loop2 is almost same at different times, the signal of the Loop1 changes. This means the perturbation of the plasma response is large under the Loop1.

In this experiment, we observe magnetic fluctuations by magnetic probes. However, the low- n MHD instability to produce the magnetic island was not observed strongly. Thus, the perturbed field may be produced by the equilibrium response.

3. Comparison of Magnetic Diagnostics and a 3D MHD Equilibrium

In this section, we study a 3D MHD equilibrium cal-

culaton and it is compared with the magnetic diagnostics.

The magnetic diagnostics observes the perturbation of the plasma response from the outside of the plasma. Observed signals are total perturbations and its depend on the internal distribution of the plasma current density. Thus, to understand the magnetic field structure, other analyses are necessary. A method is the mode analysis assuming currents filaments in the plasma. In this method, the multifilament currents $I_f = I_{mn} \cos(m\theta - n\phi + \alpha)$ are put on the resonant surface, where I_{mn} and α are the maximum filament current and the phase of the mode, respectively. Symbols θ and ϕ show poloidal and toroidal angles on the Boozer coordinate system. The subscripts m and n indicate the poloidal and toroidal mode numbers. The spatial structure of the mode is identified through the comparison between observed perturbation and the calculated flux. From these procedures, we can identify the mode and approximated location of the plasma response. Details are shown in ref. [9]. However, in this analysis, the magnetic field structure cannot be identified because filament currents are assumed. If we can get the internal distribution of the plasma current density $\mathbf{j}_{\text{plasma}}$, we can reconstruct the external perturbation from the internal $\mathbf{j}_{\text{plasma}}$ distribution. A candidate to get $\mathbf{j}_{\text{plasma}}$ is the information from the 3D MHD equilibrium calculation if the magnetic island is changed by the equilibrium response. The 3D MHD equilibrium calculation gives the information of the magnetic field ($\mathbf{j}_{\text{plasma}}$). The DIAGNO [14], V3FIT [15] and JDIA [16] codes were developed to couple a 3D MHD equilibrium code, VMEC [17]. Using those codes, we can compare observed signals and calculated perturbation without the magnetic island because the VMEC assumes perfectly nested flux surfaces. This means those code cannot identify the magnetic island breaking nested flux surfaces. To identify the magnetic island, coupling the JDIA and HINT2 codes was proposed. The HINT2 is a 3D MHD equilibrium calculation code without assumptions of perfectly nested flux surfaces [18].

In Fig. 5, flux surfaces including $n = 1$ island are shown for the vacuum and a finite- β equilibrium. The calculation was done for $R_{\text{ax}} = 3.6$ m, $A_p = 8.3$, $\kappa \sim 1$ corresponding to the experiment. These cross sections are same at the poloidal section along the Thomson scattering system (the chord is along on $Z = 0$ const. plane). The equilibrium calculation was done with the initial pressure profile $p = p_0(1 - s)(1 - s^4)$, which s is the normalized toroidal flux. A large $n = 1$ island appears for the vacuum field. However, in spite of appearing the large island, clear flux surfaces are kept in the inside and outside of $n = 1$ island. For finite- β ($\langle\beta\rangle \sim 1.5\%$), $n = 1$ island evolves and the stochasticization appears around the island. Since the plasma pressure is fluttering on the island, the Pfirsch-Schlüter (P-S) current flow is distorted. This distorted current flow generates additional perturbed field to evolve or suppress the island. In addition, the pressure-induced perturbation generates the higher-mode perturbed

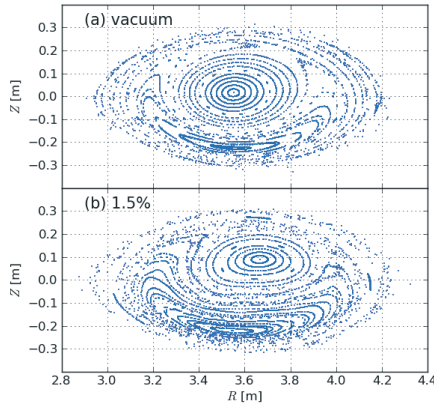


Fig. 5 Poincaré plots of magnetic field lines with $n/m = 1/1$ island are shown for the vacuum field and a finite- β equilibrium ($\beta \sim 1.5\%$). The configuration is an inward shifted configuration ($R_{ax} = 3.6$ m, $A_p = 8.3$, $\kappa \approx 1$). The color-bar indicates the connection length of magnetic field lines.

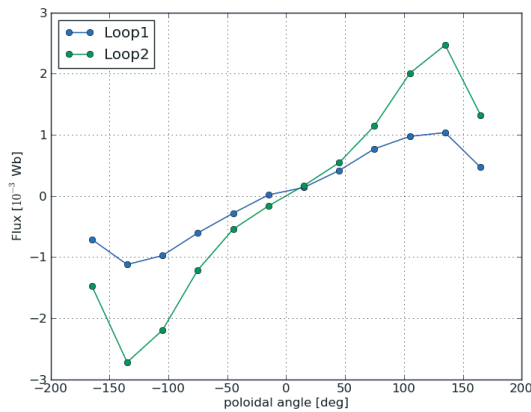


Fig. 6 Calculated flux on poloidal loop arrays are plotted corresponding to Fig. 4.

field. The nonlinear coupling of those modes leads the stochasticization of field lines.

Figure 6 shows calculated flux of poloidal loop arrays from HINT2 and JDIA corresponding to Fig. 4. The order of the flux between the calculation and observation is comparable but the factor is different. The profile is similar in both cases. Especially, the Loop1 is sensitive in the calculation and it corresponds to the experiment. From this calculation, we can guess the magnetic field structure from Fig. 5. In Fig. 3, the temperature collapse was observed but that was not the collapse. According to Fig. 5, $n/m = 1/1$ island evolved then the magnetic axis shifted upward. Therefore, the electron temperature was observed on the O-point of the magnetic island.

4. Discussion and Summary

We proposed the study of the magnetic island by coupling the magnetic diagnostics and the HINT2 code. Observed and calculated flux on poloidal loop arrays are comparable then the magnetic island evolves due to the equi-

librium response. However, the difference is found in the comparison. The order of the flux is same but the factor is a little bit different. In this study, we compare with only one calculated equilibrium assuming a pressure profile and zero net toroidal current. Since the poloidal loop array is sensitive to the pressure profile, we need to compare with other equilibria with different pressure profiles.

The net toroidal current is observed. Here, we discussed only the equilibrium response due to the distortion of the P-S current flow. If the magnetic island appears or changes, the distortion of the net toroidal currents, which are the Ohmic, beam-driven and bootstrap currents as examples, along the magnetic island is not surprising. As an example, a theory predicted the evolution and suppression of the magnetic island by the bootstrap current [19]. The consideration is analogy from the theory of the NTM. In tokamaks, the bootstrap current always destabilizes the tearing mode. However, in stellarator/heliotron, the magnetic shear is the reversed shear in tokamaks. Thus, the bootstrap current will suppress the magnetic island. For the real plasma in the experiment, the equilibrium response is coupled the effect of the P-S and other currents nonlinearly because the spontaneous evolution and suppression of the island can not be explained by only the P-S current in the experiment [5]. In that change, the magnetic island spontaneously changes due to the plasma collisionality in a discharge. To understand the equilibrium response including other currents, the extension of the HINT2 code is necessary. The HINT2 can treat only the net-toroidal current prescribed by the function of the toroidal flux. In this treatment, the vanishing of the bootstrap current in the island cannot be represented. The extension is now doing. Including the neoclassical current self-consistently, the coupling the HINT and transport codes is necessary. That is a future subject.

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- [1] H.P. Furth *et al.*, Phys. Fluids **16**, 1054 (1973).
- [2] Z. Chang *et al.*, Phys. Rev. Lett. **74**, 4633 (1995).
- [3] S. Sakakibara *et al.*, Fusion Sci. Technol. **50**, 177 (2006).
- [4] S. Sakakibara *et al.*, Fusion Sci. Technol. **58**, 176 (2010).
- [5] Y. Narushima *et al.*, Nucl. Fusion **48**, 075010 (2008).
- [6] A.H. Reiman and A.H. Boozer, Phys. Fluids **27**, 2446 (1984).
- [7] J.R. Cary and M. Kotschenreuther, Phys. Fluids **28**, 1392 (1985).
- [8] C.C. Hegna and A. Bhattacharjee, Phys. Fluids **B1**, 392 (1989).

- [9] S. Sakakibara *et al.*, Fusion Sci. Technol. **58**, 471 (2010).
[10] N. Ohyaib *et al.*, J. Nucl. Mater. **266**, 302 (1999).
[11] K. Narihara *et al.*, Phys. Rev Lett. **87**, 135002 (2001).
[12] N. Ohyaib *et al.*, Phys. Rev Lett. **88**, 055005 (2002).
[13] N. Ohyaib *et al.*, Plasma Phys. Control. Fusion **47**, 1431 (2005).
[14] H.J. Gardner, Nucl. Fusion **30**, 1417 (1990).
[15] S.P. Hirshman *et al.*, Phys. Plasmas **11**, 595 (2004).
[16] T. Yamaguchi *et al.*, Plasma Fusion Res. **1**, 011-1 (2006).
[17] S.P. Hirshman *et al.*, Comput. Phys. Commun. **43**, 143 (1986).
[18] Y. Suzuki *et al.*, Nucl. Fusion **46**, L19 (2006).
[19] C.C. Hegna and J.D. Callen, Phys. Plasmas **1**, 3135 (1994).