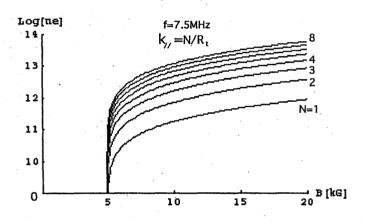
§8. RF Plasma Production on CHS by Type III Antenna

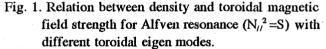
Shoji, T., Sakawa, Y. (Dept. Energy Eng. and Sci. Nagoya Univ.) Knowlton, S. (Auburn Univ.) Nishimura, K., Okamura, S.

Radio frequency waves in the ICRF range have the notable feature of producing plasmas in a wide span of magnetic field strengths, which is important in the stellarator β limits. study of The plasma density as high as <n_> =5x10¹²cm⁻³ was produced by the Nagoya Type III antenna 1-2) and this density is suitable NBI target for heating. The as а detailed structure of the plasma production magnetic behavior versus field strength has been studied in CHS. When $\omega/\omega_{ci} < 1$ and $\omega/\omega_{ci} > 1$, slow waves (shear Alfven waves) and the Ion Bernstein waves (IBW) are effectively excited by the Nagoya Type III antenna, respectively. If the Alfven resonance,

 $N_{//2}^2 = (ck_{//}/\omega)^2 = S = 1 + \omega_{pe}^2/\omega_{ce}^2 - \omega_{pi}^2/(\omega^2 - \omega_{ci}^2), (1)$ is satisfied for the slow waves, plasma is expected to be produced through the electron heating of mode converted waves resonance. at the InFig. 1. the resonance (1) is shown in density and toroidal magnetic field space, where the toroidal wave number, k ..., is determined by the toroidal eigen modes. If IBW is excited at the plasma edge where ω ≤(n+1)ω_{ci} , and propagates toward the center of the plasma ($\omega \ge n\omega_{ci}$), plasma is produced through electron Landau damping The relations between of the wave. density and toroidal magnetic field on axis, B, are shown in Fig.2. The plasma produced is above В~0.5Т and the dependence of density on B behaves like the Alfven resonance curve for the toroidal eigen mode N~4, where N is a toroidal eigen mode number. The density drop at B~0.9T in Fig.2 is not obvious in the previous experiment with a higher rf power of ~500kW. The IBW is excited the antenna where $\omega \leq 2\omega_{ci}$ near and absorbed by electrons through Landau damping around $\omega \sim \omega_{ci}$ in the plasma, which exists in the region 0.4T<B<1T. This IBW heating is confirmed by the ray trace calculations. The plasma with the mixture of H(90%) and D(10%) is examined for the IBW heating. There is 3ω_{ci} harmonics heating in the plasma when 0.4T<B<1T. The dependencies of the density on B for H and H+D plasmas are slightly different but further study is needed to explain the results. Reference

 Shoji, T., Nishimura, K, et al., Nagoya Univ. Ann.Report 6(1989)1
Nishimura, K., Shoji, T., et al., Fusion Tech. 17(1990)86





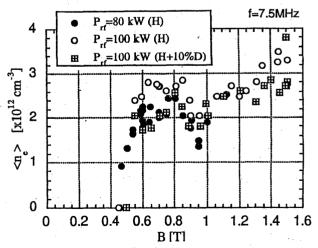


Fig. 2. Averaged plasma density as a function of B for H and H+D(10%) plasmas.